

Holographic Renormalization Group Flows

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Outlook



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Holographic picture for deviation from conformality

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- **Holographic RGF for simple models**

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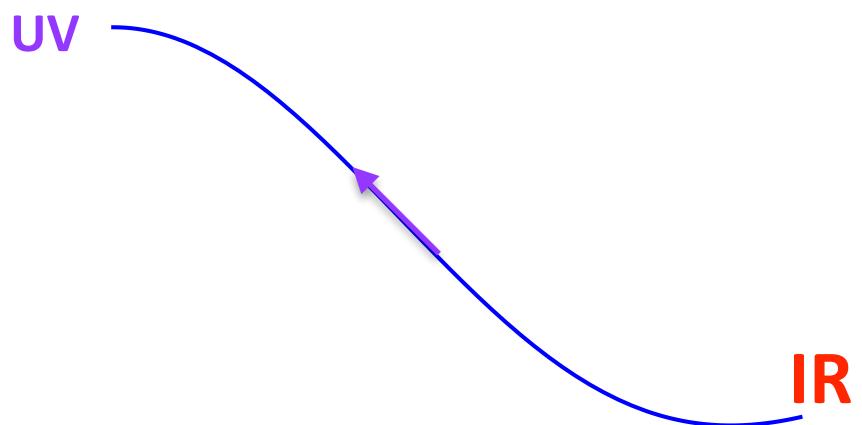
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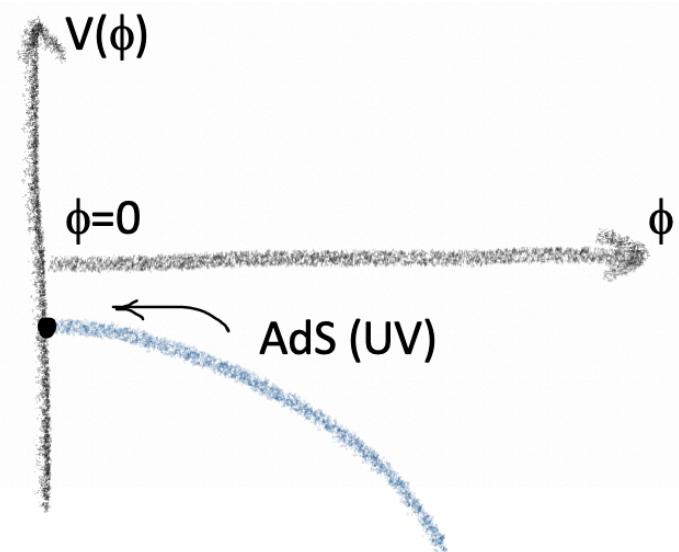
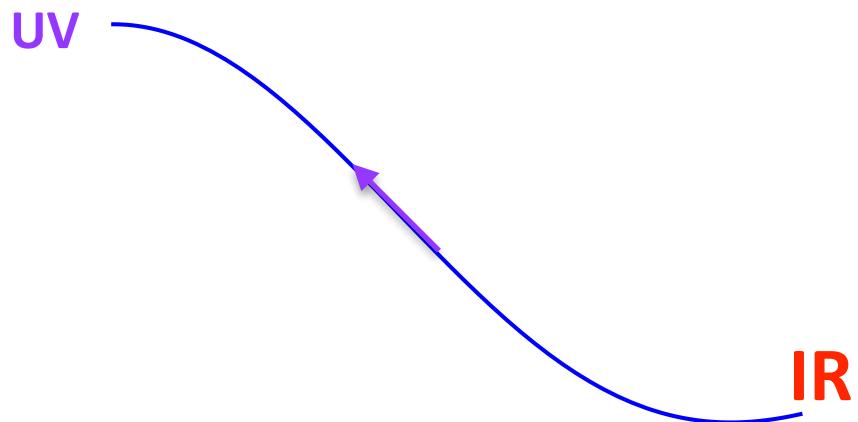
- Holographic picture for deviation from conformality
- Holographic RGF for simple models
- Exact HRGF for two exp potential
- HRGF for chemical potential $\mu \neq 0$
- HRGF for anizotropic models

Holographic picture for deviation from conformality

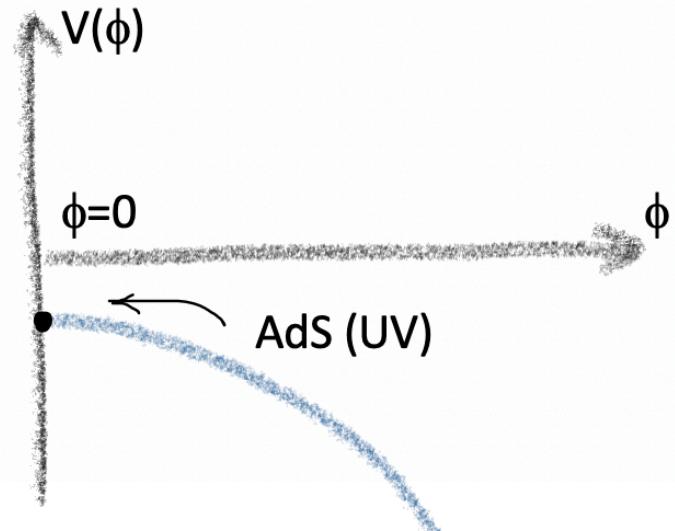
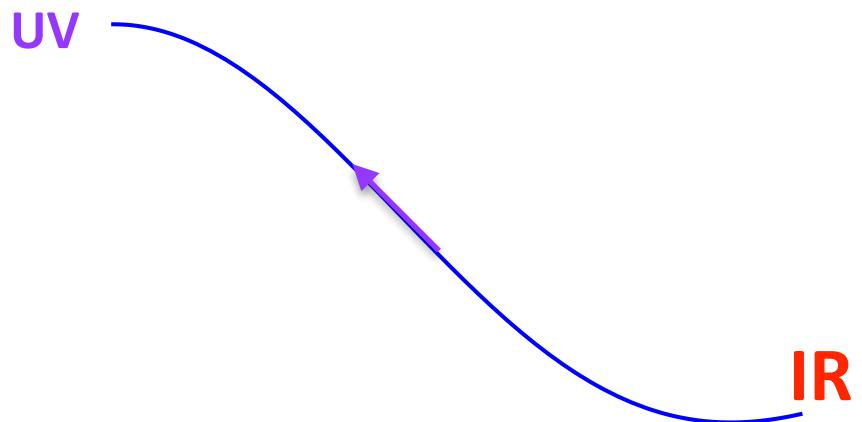
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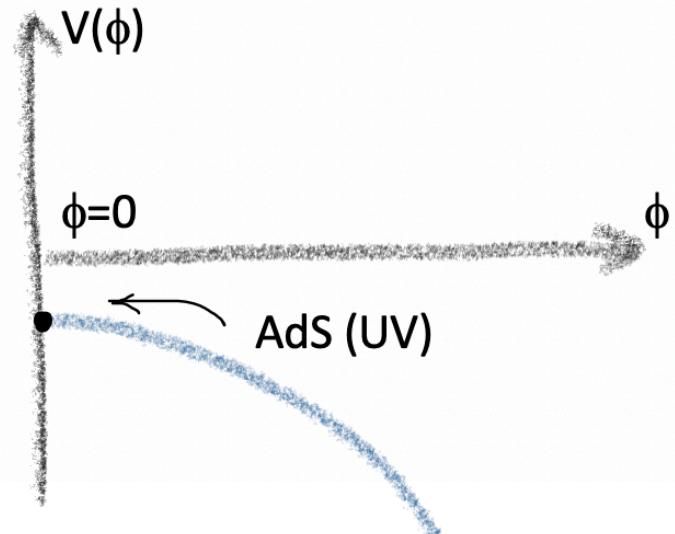
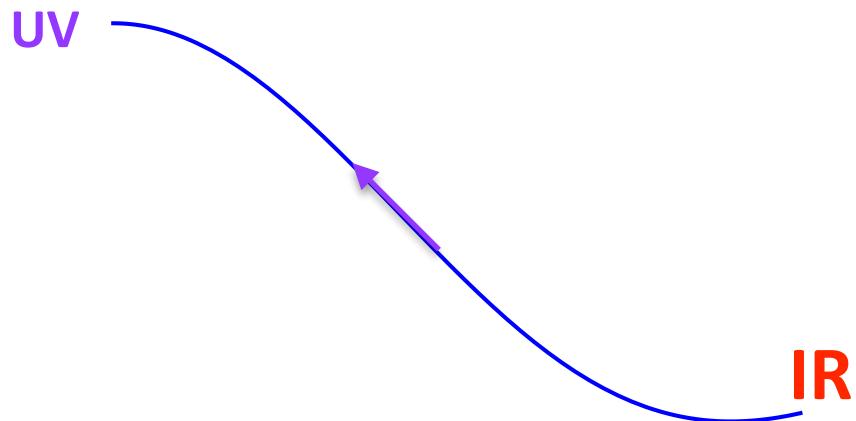
Holographic picture for deviation from conformality



- d -dim CFT has a description in terms of $d + 1$ -dim gravity in AdS :
$$S = \int dx^d du \sqrt{-g} (R - \Lambda).$$
- An operator $\mathcal{O}(x)$ corresponds to a dynamical bulk field $\phi(x, u)$
- $\phi(x, 0)$ – a source for the \mathcal{O} in the CFT

$$S = \int dx^d du \sqrt{-g} \left[R - \frac{1}{2} (\partial\phi)^2 - V(\phi) \right].$$

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- $\phi(x, u) = \alpha u^{d-\Delta} + \dots \Leftrightarrow S = S_{CFT} + \int d^4 x \alpha \mathcal{O}(x)$
- $\alpha = 0$ – undeformed CFT, bulk scalar – const., spacetime is AdS
- $\alpha \neq 0$ corresponds to relevant coupling for the CFT; deform. AdS

Holographic Renormalization Group

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The domain wall solution (d -dimensional Poincaré invariant)

$$ds^2 = e^{2\mathcal{A}(u)} \eta_{ij} dx^i dx^j + du^2, \quad \phi = \phi(u)$$

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- The β -function

$$\beta = \frac{d\lambda}{d \log E} = \frac{d\phi}{d\mathcal{A}}$$

Starting point - 5-dim background

Gravity + scalar field

Gravity + scalar field

$$S = \int \frac{d^5x}{16\pi G_5} \sqrt{-\det(g_{\mu\nu})} \left[R - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - V(\phi) \right]$$

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$$ds^2 = B^2(z) [-dt^2 + dx^2 + dy_1^2 + dy_2^2 + dz^2], \quad \phi = \phi(z)$$

EOM:

$$\frac{\ddot{B}}{B} + \frac{2\dot{B}^2}{B^2} + \frac{3\dot{B}\dot{g}}{2B} + \frac{B^2V}{3} = 0,$$

$$\ddot{\phi} + \dot{\phi} \left(\frac{3\dot{B}}{B} \right) - B^2 \partial_\phi V = 0,$$

Const.

$$\frac{\ddot{B}}{B} - \frac{2\dot{B}^2}{B^2} + \frac{\dot{\phi}^2}{6} = 0$$

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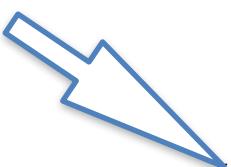
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RGF equation:

$$\frac{dX}{d\phi} = -\frac{4}{3} \left(1 - \frac{3}{8} X^2\right) \left(1 + \frac{1}{X} \frac{\partial_\phi V}{V}\right)$$

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Compare with the method of super potential:

$$V = \frac{1}{8}W'(\phi)^2 - \frac{1}{3}W(\phi)^2$$

Gravity + scalar field. Details, 1/2

$$\dot{X} = \frac{1}{3} \left\{ \frac{B}{\dot{B}} \left(-\dot{\phi} \right) \left(\frac{\dot{g}}{g} + \frac{3\dot{B}}{B} \right) + \frac{B^3 \partial_\phi V}{g \dot{B}} - \frac{\dot{\phi} B^2}{\dot{B}^2} \left(\frac{2\dot{B}^2}{B^2} - \frac{\dot{\phi}^2}{6} \right) + \dot{\phi} \right\}. \quad (*)$$

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Dividing on $\dot{\phi}$ gives

$$\frac{3\dot{X}}{\dot{\phi}} = -4 - \frac{B\dot{g}}{\dot{B}g} + \frac{B^3 \partial_\phi V}{g\dot{\phi}\dot{B}} + \frac{3}{2} X^2. \quad (\#)$$

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Now our goal is to exclude \dot{B} and $\dot{\phi}$ from the RHS of $(\#)$. We do this in 3 steps.

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$$\frac{3\dot{X}}{\dot{\phi}} = -4 + \frac{3XB^2\partial_\phi V}{g\dot{\phi}^2} + \frac{3}{2} X^2. \quad (**)$$

Gravity + scalar field. Details, 2/2

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Step 2

Gravity + scalar field. Details, 2/2

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Step 2

- We get relation for $\dot{\phi}^2$ and substitute it into (**).

$$\frac{4\dot{\phi}^2}{9X^2} + \frac{B^2V}{3g} - \frac{\dot{\phi}^2}{6} = 0. \quad (***)$$

We rewrite (***) as

$$\left(\frac{1}{6} - \frac{4}{9X^2}\right)\dot{\phi}^2 = \frac{2V}{6g}B^2 \quad (***)$$

and from (****) get combination $g\dot{\phi}^2/B^2$

$$\frac{g\dot{\phi}^2}{B^2} = \frac{2V}{1 - \frac{8}{3X^2}}. \quad (*****)$$

Gravity + scalar field. Details, 2/2

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$$\left(\frac{1}{6} - \frac{4}{9X^2}\right)\dot{\phi}^2 = \frac{2V}{6g}B^2 \quad (***)$$

and from (****) get combination $g\dot{\phi}^2/B^2$

$$\frac{g\dot{\phi}^2}{B^2} = \frac{2V}{1 - \frac{8}{3X^2}}. \quad (*****)$$

Step 3

Substituting in the RHS of (**) the expression (*****) we finally get

$$\frac{3\dot{X}}{\dot{\phi}} = -4 \left(1 - \frac{3}{8}X^2\right) \left(1 + \frac{1}{X} \frac{2\partial_\phi V}{2V}\right).$$

Gravity + scalar field. Simple examples

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Conformal case. $V=\text{const}$

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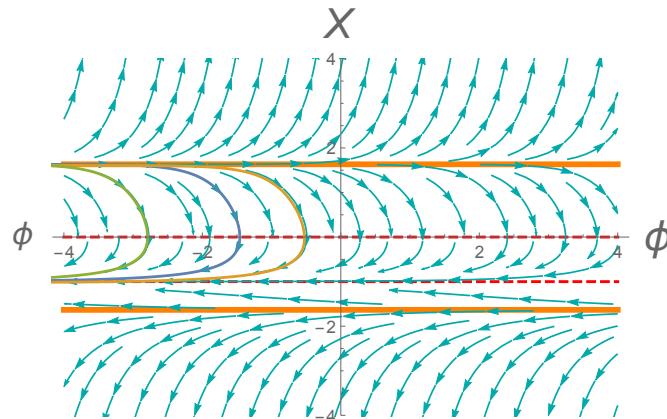
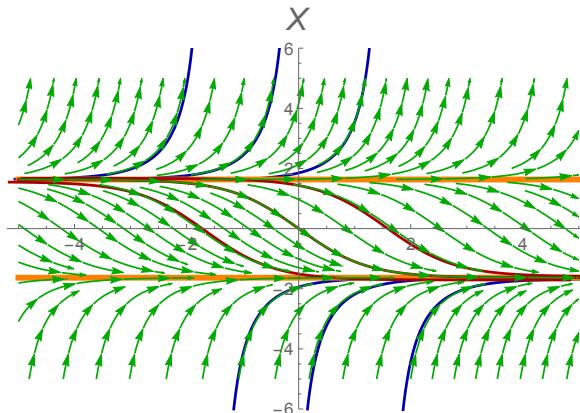
Dilaton gravity with one exponential potential

Chamblin Reall model $V(\phi) = \exp(c\phi) + V_0$

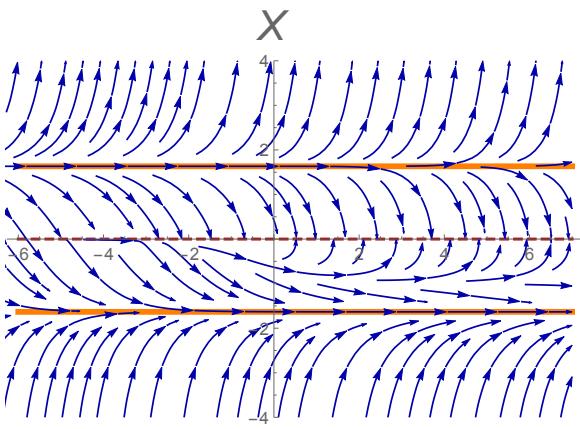
$$c = 1$$

$$\phi = \phi_0 + \frac{1}{5} \left[\left(2\sqrt{6} - 3 \right) \log |2\sqrt{6} - 3X| + 6 \log |X + 1| - \left(3 + 2\sqrt{6} \right) \log |3X + 2\sqrt{6}| \right]$$

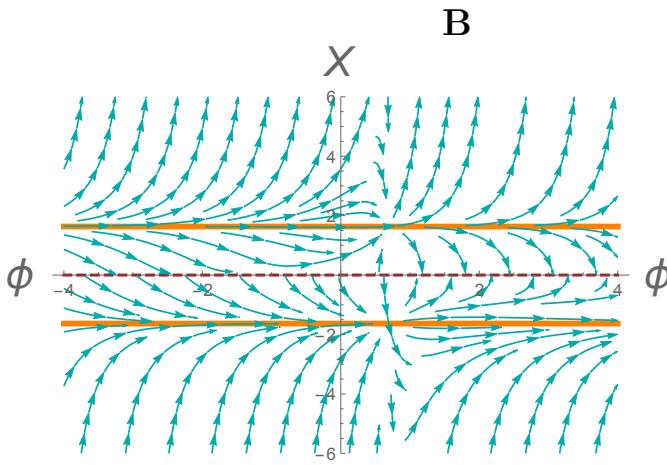
Stream-plots for vectors field



A



C



D

- A) AdS, i.e. $V(\phi) = \text{const}$, B) $V(\phi) = \exp(c\phi)$, C) $V(\phi) = \exp(c\phi) + V_0$, $V_0 > 0$
D) $V(\phi) = \exp(c\phi) + V_0$, $V_0 < 0$.

Holographic RG Flow

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$$V_{AGP}(\varphi) = C_1 e^{2k\varphi} + C_2 e^{\frac{32}{9k}\varphi}$$

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IA, A.Golubtsova,
G.Policastro,
1803.06764

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IA, A.Golubtsova,
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1803.06764

Generalization of Chamblin&Reall model,
hep/th. 9903225

Two exponent potential

Two exponent potential

$$V = C_1 e^{2\mathbf{k}\phi} + C_2 e^{\frac{32}{9\mathbf{k}}\phi}; \quad C_1 < 0, C_2 > 0$$

Two exponent potential

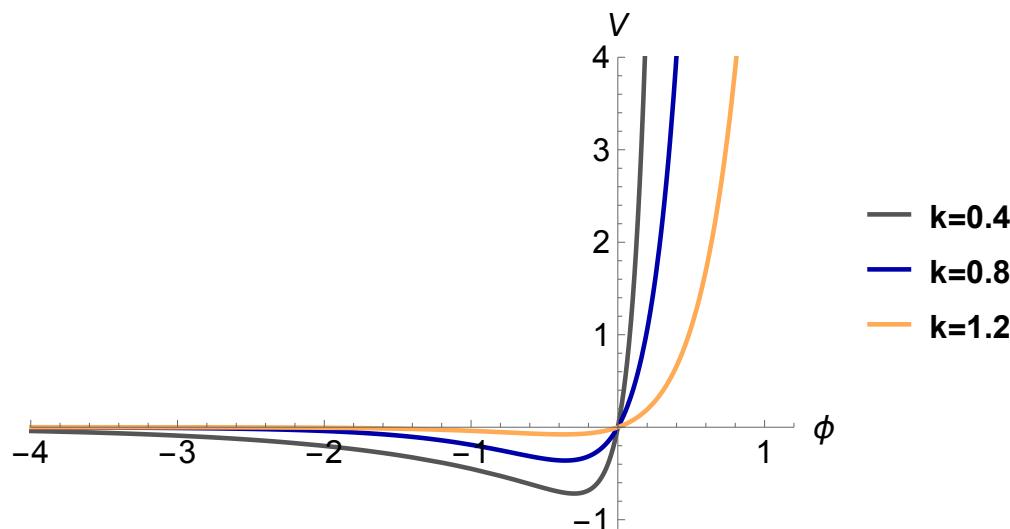
IA, A.Golubtsova,
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1803.06764

$$V = C_1 e^{2\kappa\phi} + C_2 e^{\frac{32}{9\kappa}\phi}; \quad C_1 < 0, C_2 > 0$$

Two exponent potential

IA, A.Golubtsova,
G.Policastro,
1803.06764

$$V = C_1 e^{2k\phi} + C_2 e^{\frac{32}{9k}\phi}; \quad C_1 < 0, C_2 > 0$$



Exact solutions for the metric and the dilaton (zero temperature)

$$ds^2 = F_1^{\frac{8}{9k^2-16}} F_2^{\frac{9k^2}{2(16-9k^2)}} (-dt^2 + d\vec{y}^2) + F_1^{\frac{32}{9k^2-16}} F_2^{\frac{18k^2}{16-9k^2}} du^2$$

$$\phi = -\frac{9k}{9k^2-16} \log F_1 + \frac{9k}{9k^2-16} \log F_2$$

Two exponent potential

Two exponent potential

- $F_1 = \sqrt{\left| \frac{C_1}{2E} \right|} \sinh(\mu |u - u_{01}|), \mu = \sqrt{\left| \frac{3E}{2} \left(k^2 - \frac{16}{9} \right) \right|},$
 $F_2 = \sqrt{\left| \frac{C_2}{2E} \right|} \sinh\left(\frac{4}{3k} \mu |u - u_{02}| \right).$
- $F_1 = \sqrt{\frac{3}{4} \left(k^2 - \frac{16}{9} \right) C_1} (u - u_{01}), F_2 = \sqrt{\frac{4}{3} \left(\frac{16}{9k^2} - 1 \right) C_2} |u - u_{02}|$
 $E = 0$

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E, u_{01}, u_{02} are constants of integration

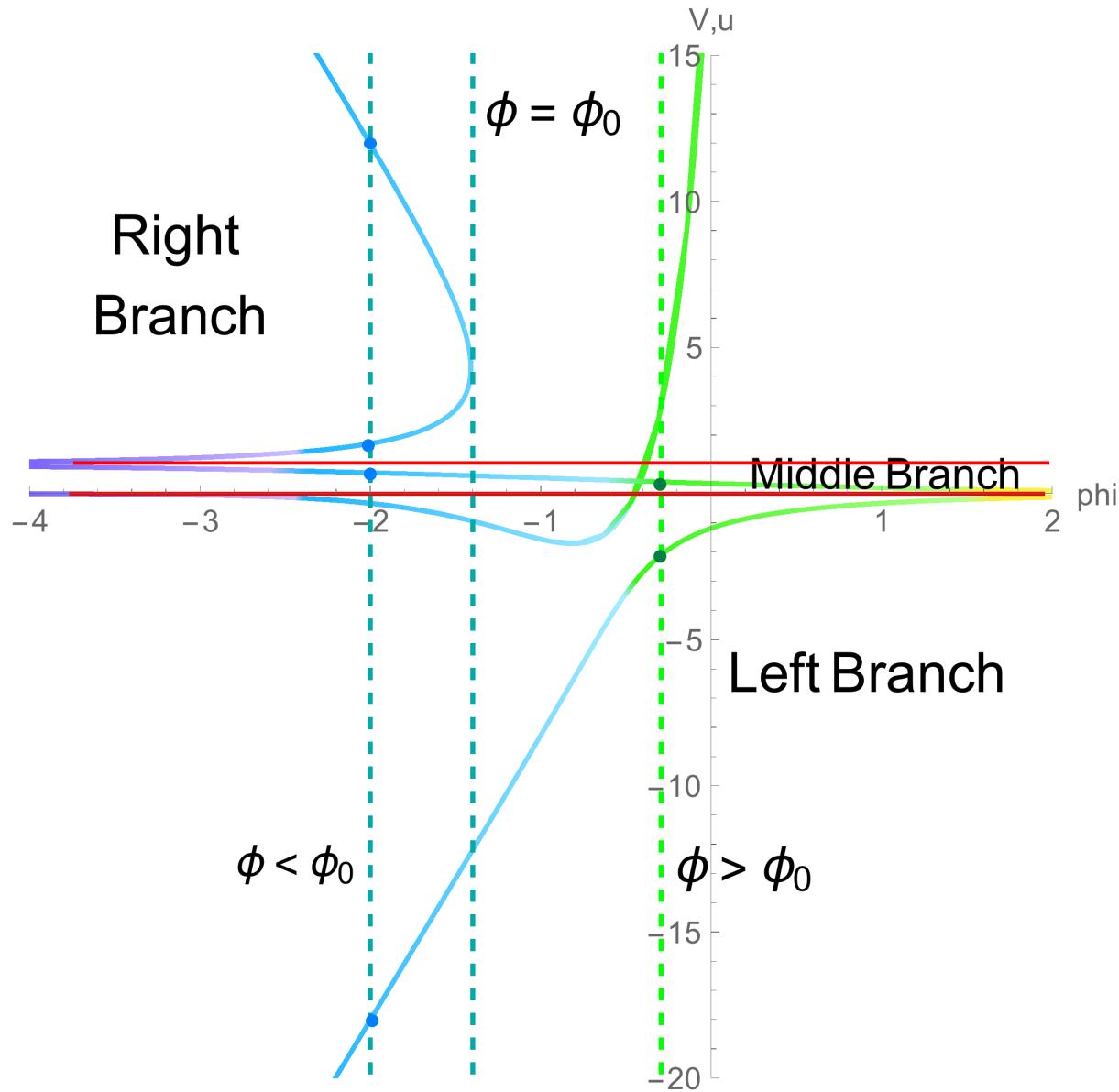
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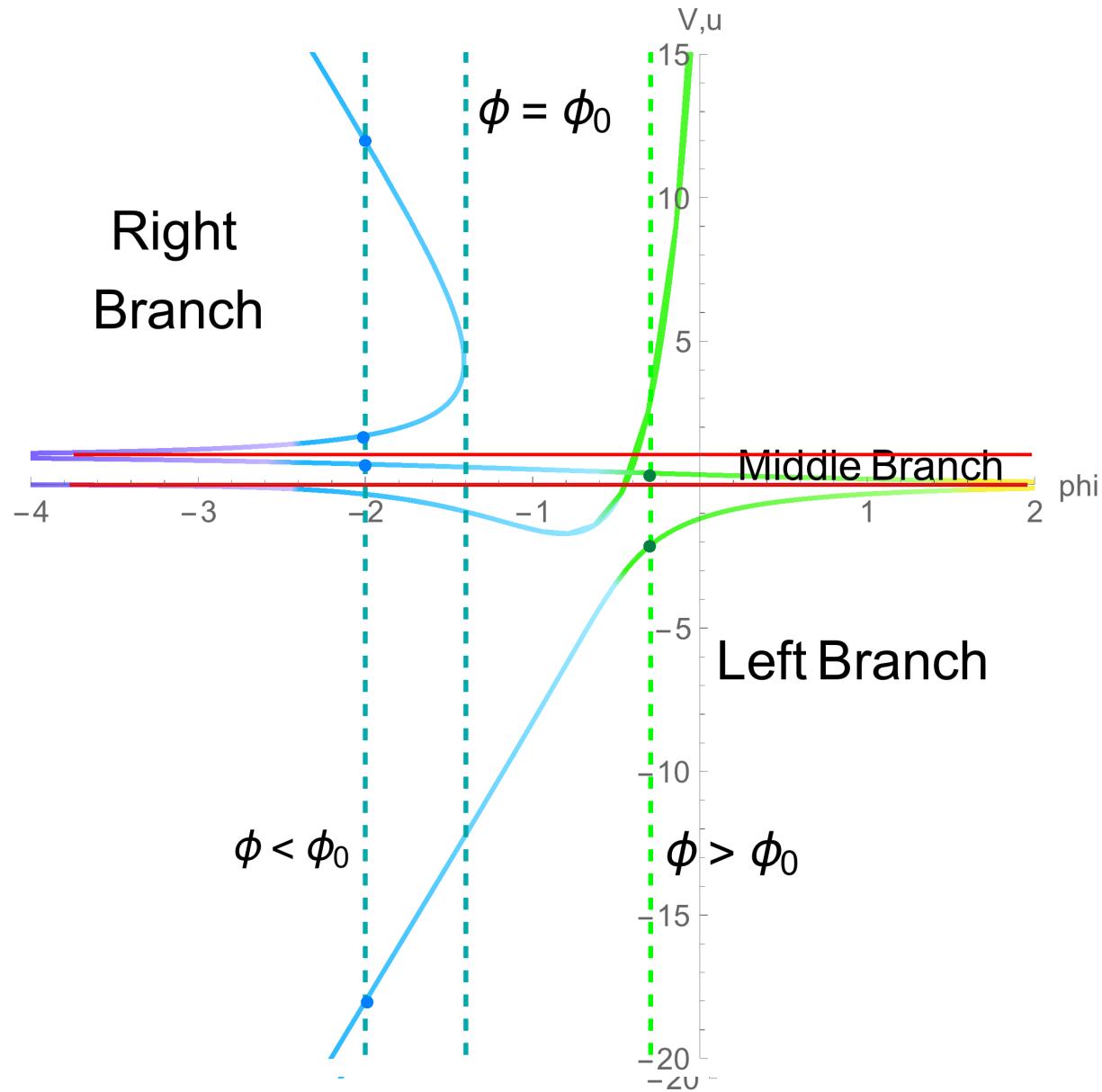
E, u_{01}, u_{02} are constants of integration

- types of RG flows: **1)** on $u \in (u_{02}, u_{01})$, **2)** $u \in (u_{01}, +\infty)$.
- boundaries of the backgrounds correspond to fixed points.
- $u_{01} = u_{02} \Rightarrow$ special RG flow : $u \in (u_0, +\infty)$ AdS **UV** fixed point

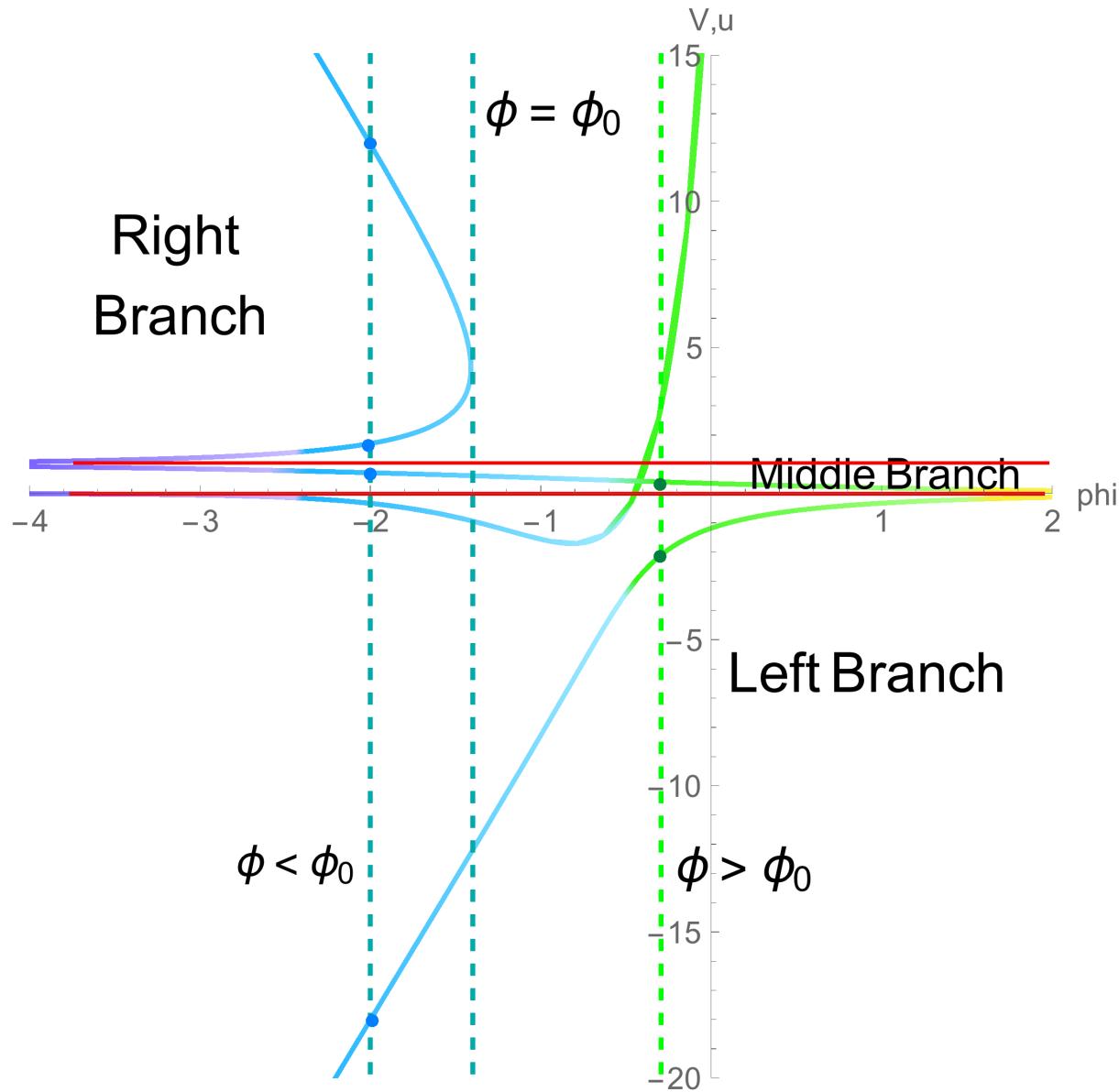
The solution for the dilaton - 3 regions!



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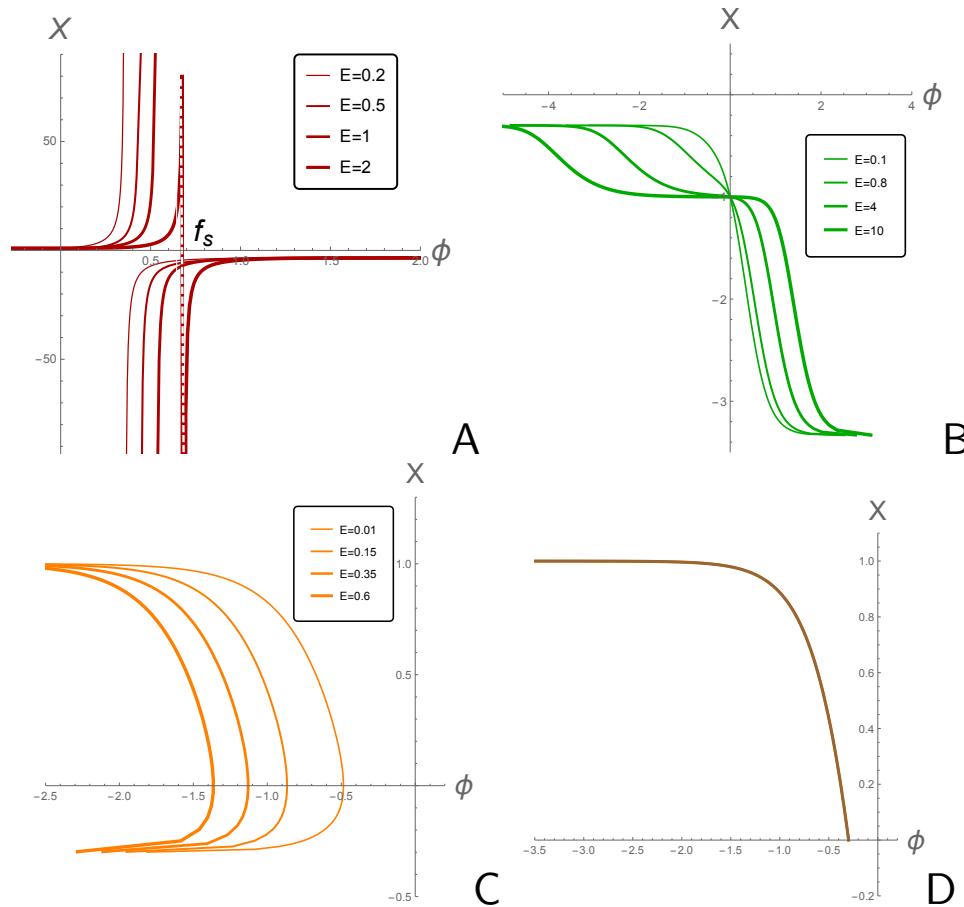
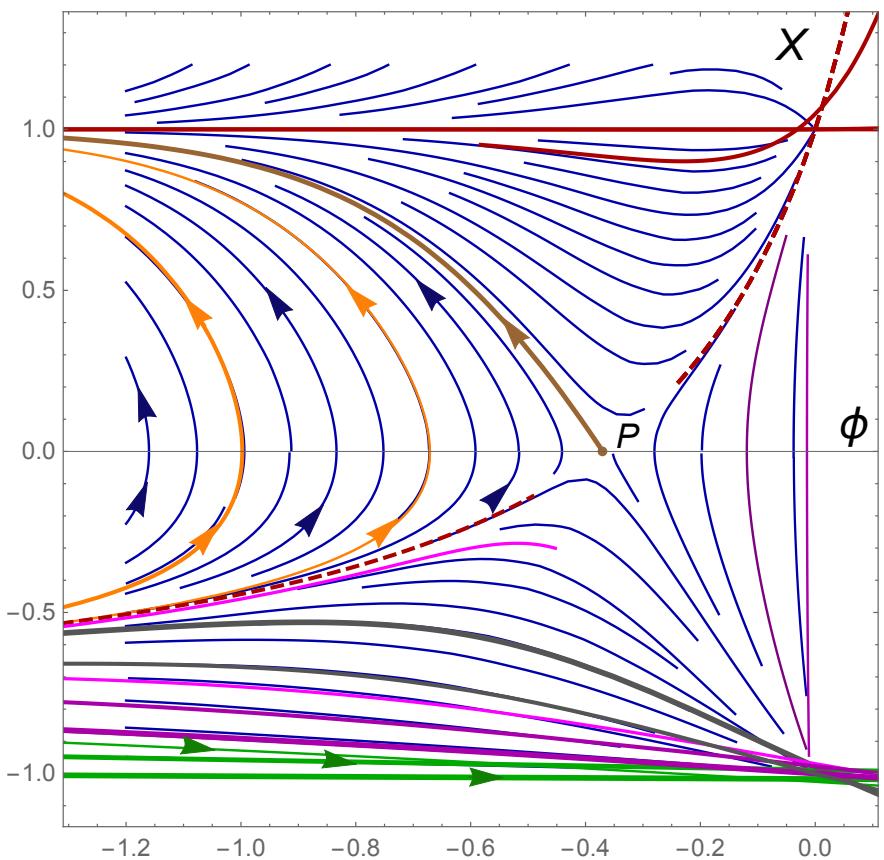
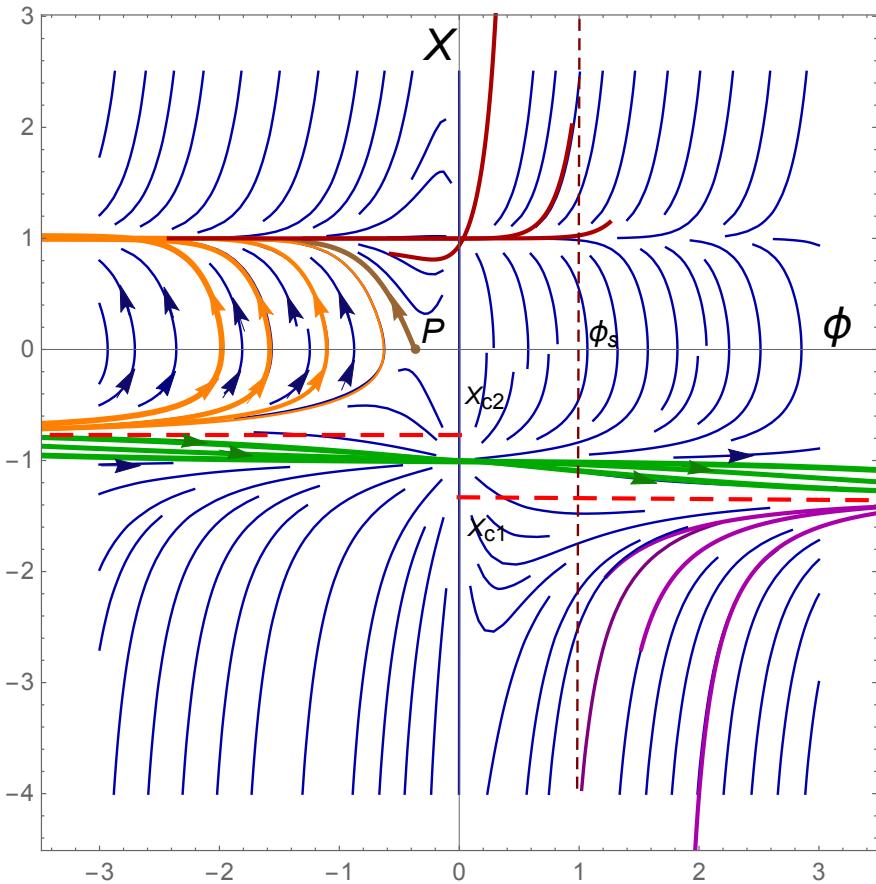


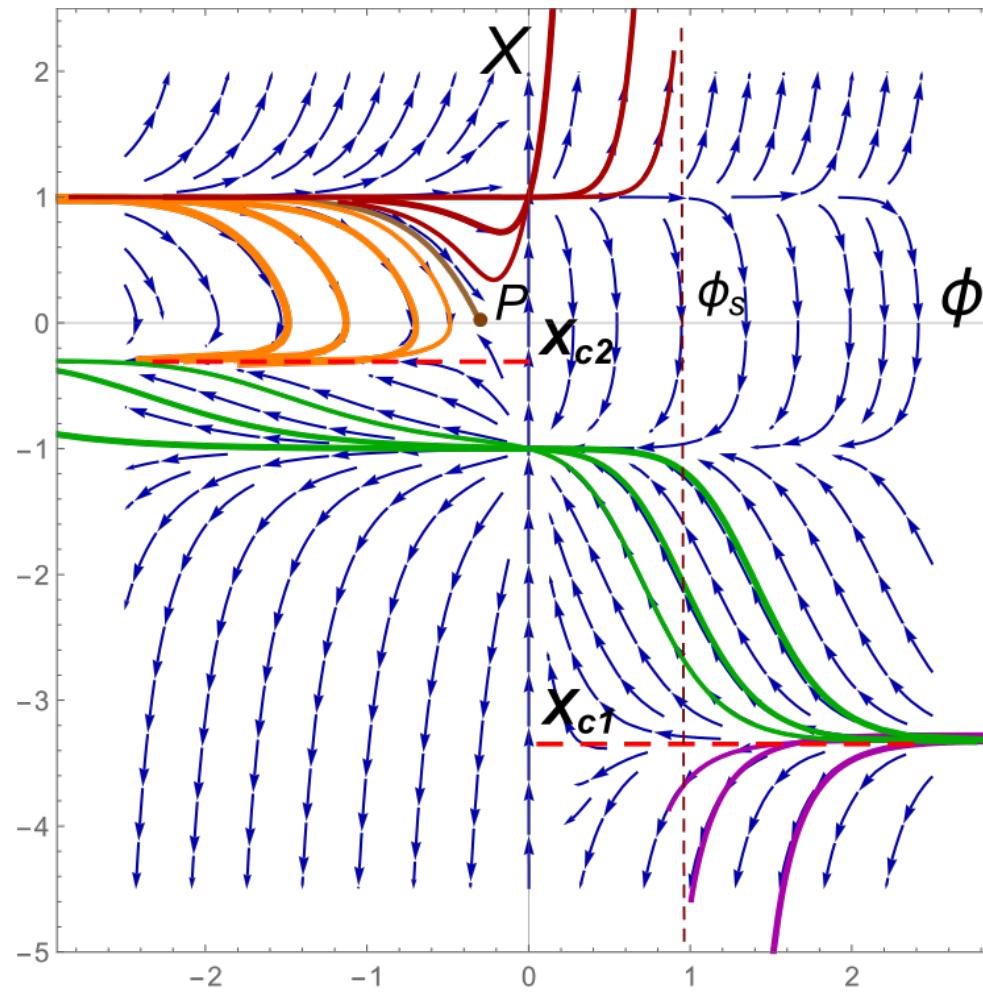
Figure: The behaviour of the X -function with the dependence on the dilaton plotted using the solutions for \mathcal{A} . A) left B) middle C) right D) $u_{01} = u_{02}$

$$\frac{dX}{d\phi}=-\frac{4}{3}\left(1-X^2\right)\left(1+\frac{3}{8X}\frac{d\log V}{d\phi}\right)$$

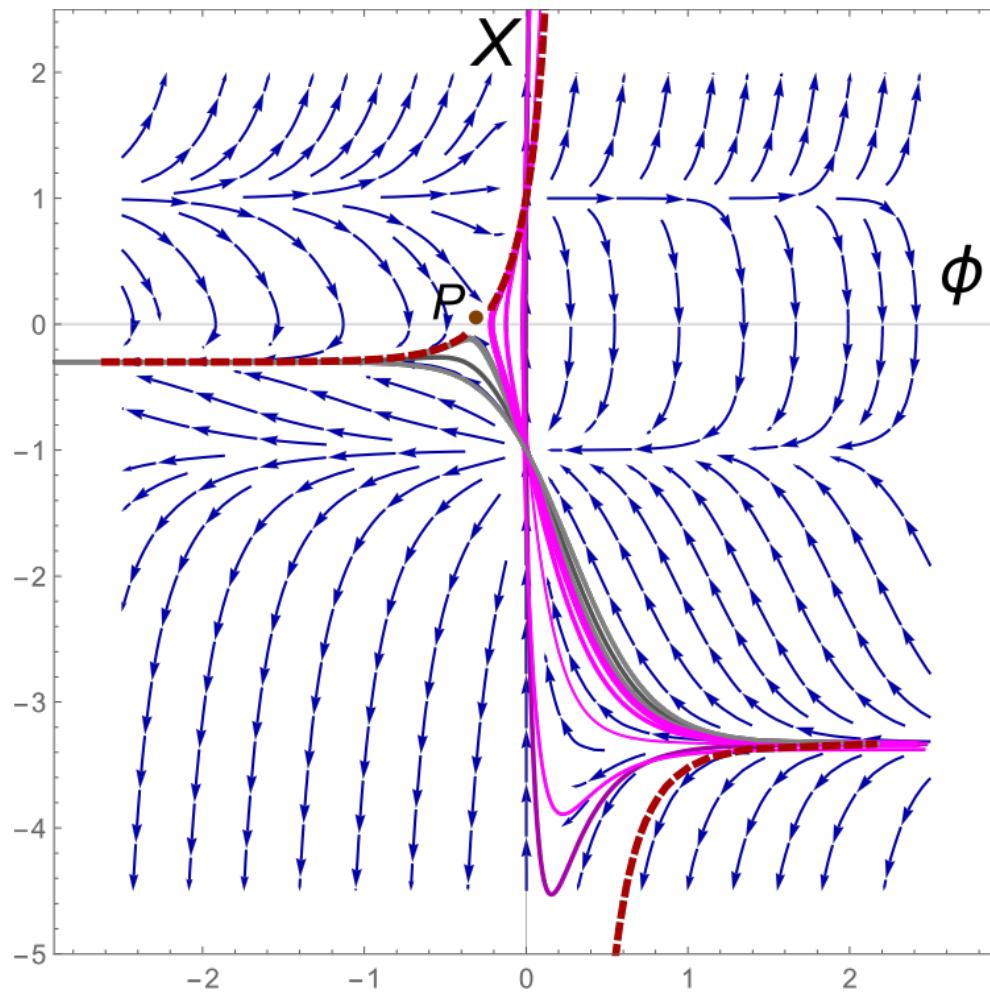
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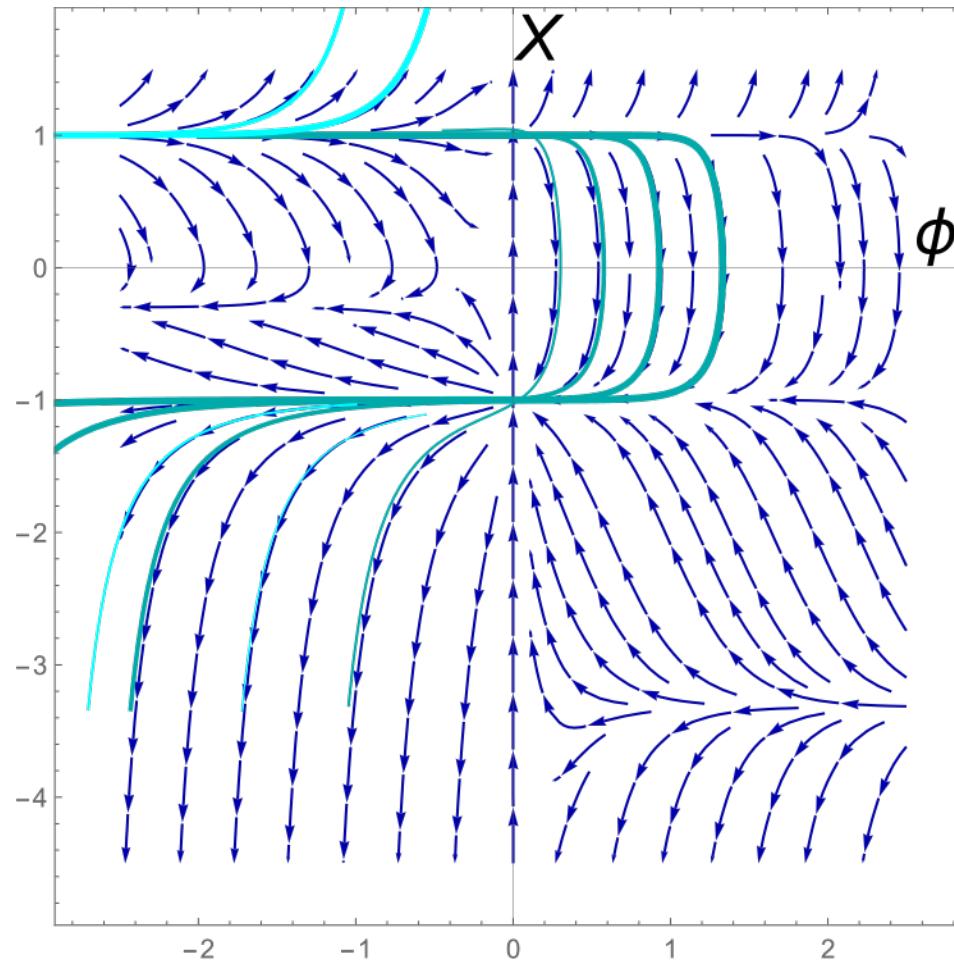
\sinh – solutions



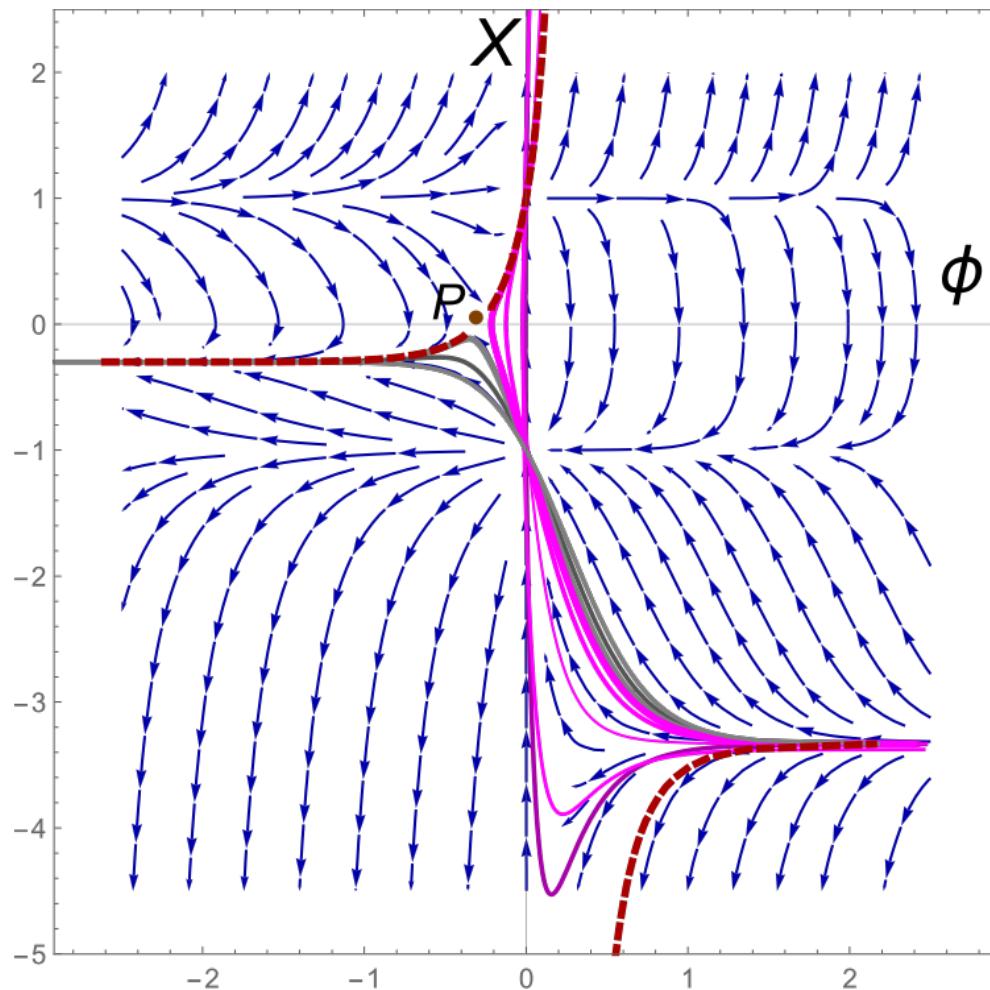
$\sin - solutions$



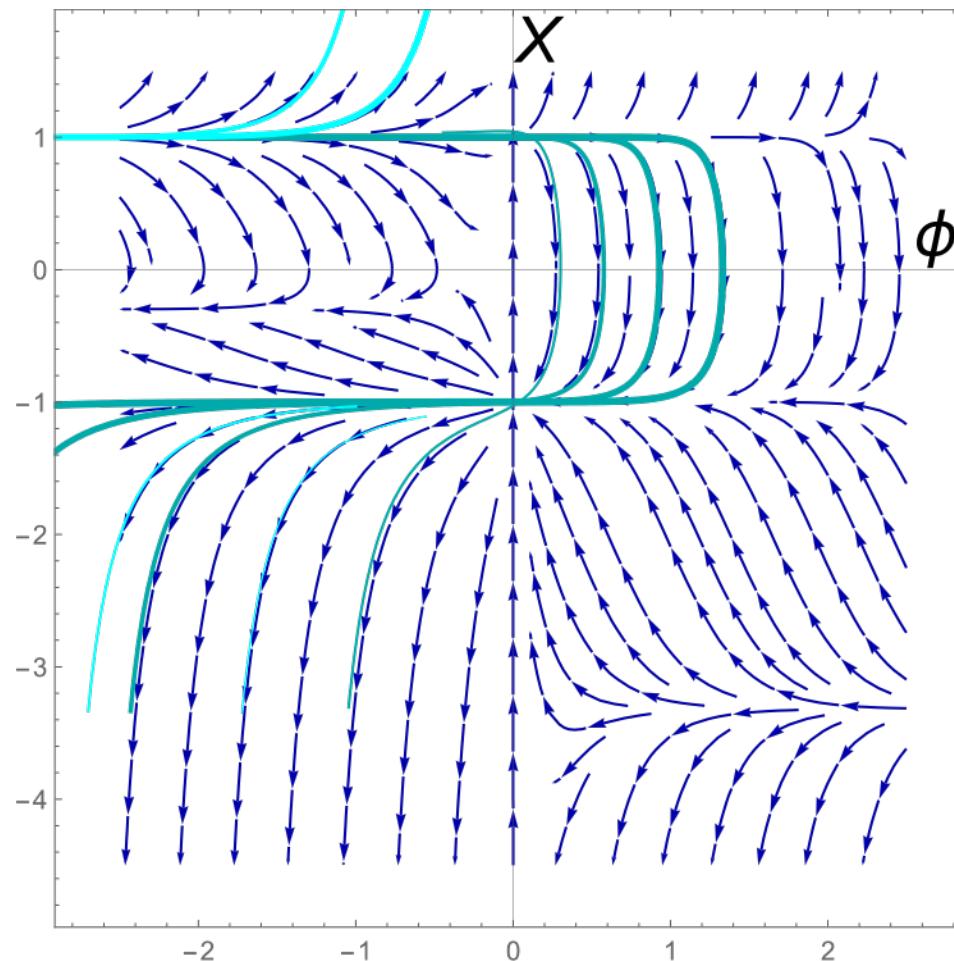
\cosh – solutions



$\sin - solutions$



\cosh – solutions



EOM as an autonomous system

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IA, A.Golubtsova, A. G Policastro,

M Skugoreva, Work in progress

EOM as an autonomous system

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The domain wall ansatz

M Skugoreva, Work in progress

$$ds^2 = e^{2A(w)}(-dt^2 + d\vec{x}^2) + dw^2, \quad \phi = \phi(w)$$

Introduce new variables

$$X = \frac{1}{3} \frac{\dot{\phi}}{\dot{A}}, \quad Z = \frac{1}{e^{\frac{2(16-9k^2)}{9k}\phi} + 1},$$

$$V = C_1 \left(\frac{1-Z}{Z} \right)^{\frac{9k^2}{16-9k^2}} + C_2 \left(\frac{1-Z}{Z} \right)^{\frac{16}{16-9k^2}}.$$

Then zeros of the potential are $Z = \frac{C_2}{C_2 - C_1}$.

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Then zeros of the potential are $Z = \frac{C_2}{C_2 - C_1}$.

$$\frac{dZ}{dA} = \frac{2(9k^2 - 16)}{3k}(1 - Z)ZX,$$

$$\frac{dX}{dA} = (X^2 - 1) \left(4X + \frac{16C_2 + Z(C_1 9k^2 - 16C_2)}{3k(C_2 + Z(C_1 - C_2))} \right)$$

Autonomous system for holographic RG flows at $T = 0$

IA, AG, GP, MS'20

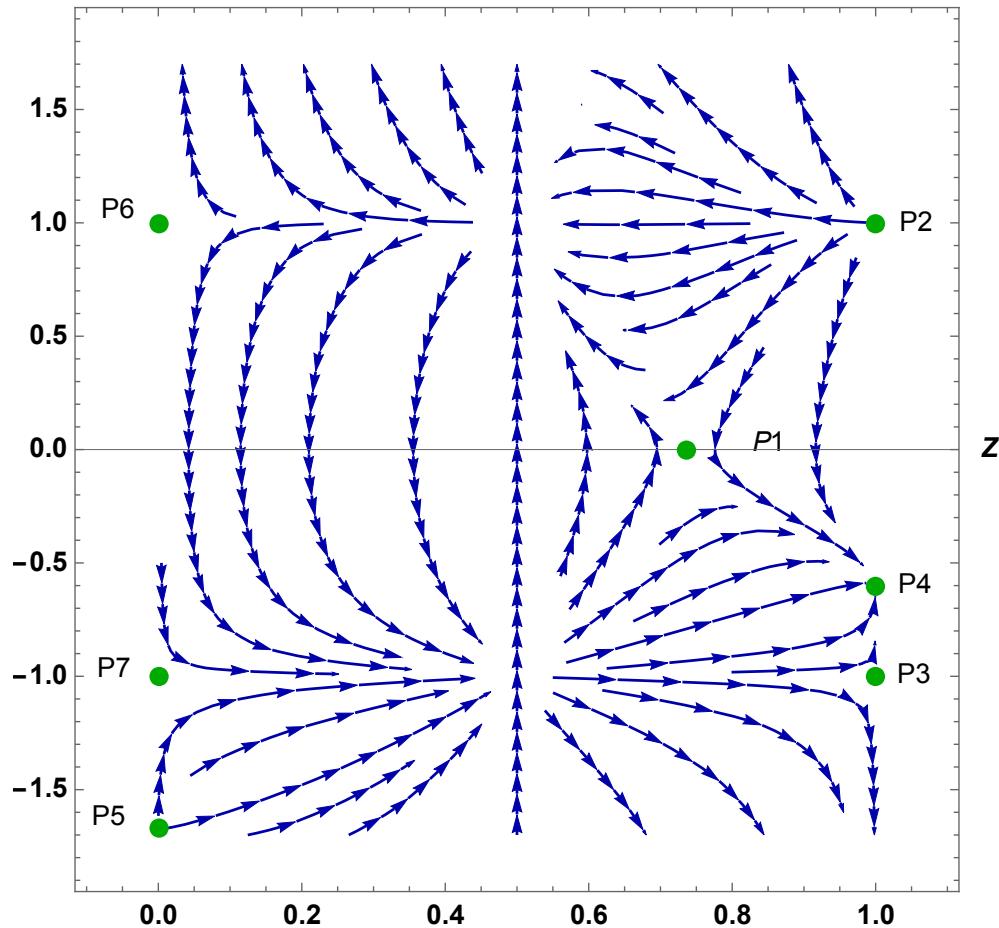


FIGURE: Phase portrait for the system at $T = 0$ with $k = 0.8$, the direction of arrows show with respect to increasing A

Autonomous system for holographic RG flows at $T = 0$

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To classify stability of equilibrium points we do the linear perturbation δZ and δX around each from the found critical points (Z_c, X_c) in order

$$Z = Z_c + \delta Z, \quad X = X_c + \delta X.$$

Substituting these expressions into the equations of motion we get

the linearized equations of motion:

Autonomous system for holographic RG flows at $T = 0$

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$$\frac{d}{dA} \begin{pmatrix} \delta Z \\ \delta X \end{pmatrix} = \mathcal{M} \begin{pmatrix} \delta Z \\ \delta X \end{pmatrix},$$

where

$$\mathcal{M} = \left(\begin{array}{cc} \frac{\partial f}{\partial Z} & \frac{\partial f}{\partial X} \\ \frac{\partial g}{\partial Z} & \frac{\partial g}{\partial X} \end{array} \right) \Bigg|_{Z=Z_c, X=X_c}$$

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is the perturbation matrix. The elements of Jacobian matrix \mathcal{M} are

$$m_{11} = \frac{2(9k^2 - 16)}{3k}(1 - 2Z_c)X_c, \quad m_{12} = \frac{2(9k^2 - 16)}{3k}(1 - Z_c)Z_c$$

$$m_{21} = \frac{C_1 C_2 (X_c^2 - 1)(9k^2 - 16)}{3k(C_2 + Z_c(C_1 - C_2))^2},$$

$$m_{22} = 4(3X_c^2 - 1) + 2X_c \left[\frac{16C_2 + Z_c(C_1 9k^2 - 16C_2)}{3k(C_2 + Z_c(C_1 - C_2))} \right].$$

Backup

Backup

1. $X = 0$, $Z = \frac{16C_2}{16C_2 - 9k^2 C_1}$. $V = \text{const.}$ The eigenvalues $\lambda_1 = 4$, $\lambda_2 = -8$, a saddle. The scalar field and the metric

$$\begin{aligned}\phi_{P_1} &= \frac{9k}{2(16 - 9k^2)} \ln \left(-\frac{9k^2}{16} \frac{C_1}{C_2} \right), \quad \frac{C_1}{C_2} < 0, \\ ds^2 &= e^{2\mathcal{C}(w-w_0)} (-dt^2 + d\vec{x}^2) + dw^2,\end{aligned}$$

where w_0 is a constant of integration, $\mathcal{C} = \pm \sqrt{\frac{-C_1(16-9k^2)}{192}} \left(-\frac{9k^2}{16} \frac{C_1}{C_2} \right)^{\frac{9k^2}{2(16-9k^2)}}$

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2. $X = 1, Z = 1$. $V \rightarrow 0$. The eigenvalues $\lambda_{1,2} = \frac{2}{3k}(4 + 3k)(2 \pm |2 - 3k|)$. An unstable node. The scalar field and the metric

$$\begin{aligned}\phi(w) &= \frac{3}{4} \ln \left| \frac{w - w_0}{w_2} \right| \rightarrow -\infty \quad \text{for } w \rightarrow w_0, \\ ds^2 &= \left| \frac{w - w_0}{w_1} \right|^{1/2} (-dt^2 + d\vec{x}^2) + dw^2.\end{aligned}$$

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3. $X = -1, Z = 1$. $V \rightarrow 0, \lambda_1 = 2(4 - 3k) > 0, \lambda_2 = \frac{2}{3k}(9k^2 - 16) < 0$ – a saddle.

$$\phi(w) = -\frac{3}{4} \ln \left| \frac{w}{w_2} \right| \rightarrow -\infty \text{ for } w \rightarrow +\infty, \quad ds^2 = \left| \frac{w}{w_1} \right|^{1/2} (-dt^2 + d\vec{x}^2) + dw^2,$$

where w_1, w_2 are constants.

Backup

4. $X = -\frac{3k}{4}$, $Z = 1$, $V \rightarrow 0$. A stable node: $\lambda_1 = \frac{1}{4}(9k^2 - 16)$, $\lambda_2 = \frac{1}{2}(9k^2 - 16)$

$$\phi(w) = -\frac{1}{k} \ln \left| \frac{w}{w_2} \right| \rightarrow -\infty \text{ for } w \rightarrow +\infty, \quad ds^2 = \left| \frac{w}{w_1} \right|^{\frac{8}{9k^2}} (-dt^2 + d\vec{x}^2) + dw^2,$$

5. $X = -\frac{4}{3k}$, $Z = 0$. $\lambda_1 = \frac{8}{9k^2}(16 - 9k^2)$, $\lambda_2 = \frac{4}{9k^2}(16 - 9k^2)$. An unstable node. $V \rightarrow +\infty$.

$$\phi = -\frac{9k}{16} \ln \left| \frac{w - w_0}{w_2} \right| \rightarrow +\infty \text{ for } w \rightarrow w_0, \quad ds^2 = \left| \frac{w - w_0}{w_1} \right|^{\frac{9k^2}{32}} (-dt^2 + d\vec{x}^2) + dw^2$$

where w_0 , w_1 and w_2 are some constants of integration.

6. $X = 1$, $Z = 0$, $V \rightarrow +\infty$. $\lambda_1 = \frac{8}{3k}(3k + 4) > 0$, $\lambda_2 = \frac{2}{3k}(9k^2 - 16) < 0$, a saddle.
The asymptotic form of the metric and the scalar field

$$\phi = \frac{3}{4} \ln \left| \frac{w}{w_2} \right| \rightarrow +\infty \text{ for } w \rightarrow +\infty, \quad ds^2 = \left| \frac{w}{w_1} \right|^{\frac{1}{2}} (-dt^2 + d\vec{x}^2) + dw^2.$$

where w_1 and w_2 are some constants of integration.

7. $X = -1$, $Z = 0$. $\lambda_1 = \frac{2}{3k}(16 - 9k^2) > 0$, $\lambda_2 = \frac{8}{3k}(3k - 4) < 0$, a saddle.
 $V(\phi) \rightarrow +\infty$.

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Backup

Types of fixed points

P_1	P_2	P_3	P_4	P_5	P_6	P_7
$V = const$	$V \rightarrow 0$	$V \rightarrow 0$	$V \rightarrow 0$	$V \rightarrow \infty$	\times	\times
ustable	stable	unstable	ustable	stable	\times	\times
UV	IR	UV	UV	IR	\times	\times

(We take into account that A should have the opposite direction)

Examples of flows:

- $P_1(AdS -UV)-P_2$ (hyperscaling vioation in IR)
- P_4 (hypersc. v. in UV)- P_2 (hypersc. v. in IR), bouncing solution AGP'18
- P_4 (hypersc. v. in UV)- P_5 (hypersc. v. in IR) confining solution
(AG,Ngyuen Vu'1906.12316)

Black holes in terms of w-coordinates

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$$ds^2 = e^{2A(w)}(-f(w)dt^2 + d\vec{x}^2) + \frac{dw^2}{f(w)},$$

and the dilaton is

$$\lambda = e^{\phi(w)}.$$

The function f is a so-called blackening function. It is convenient to introduce a new variable

$$g = \ln f.$$

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The function f is a so-called blackening function. It is convenient to introduce a new variable

$$g = \ln f.$$

The equations of motion with respect to w -variable are

$$12\dot{A}^2 + 3\dot{A}\dot{g} - \frac{4}{3}\dot{\phi}^2 + e^{-g}V = 0,$$

$$\ddot{A} + \frac{4}{9}\dot{\phi}^2 = 0,$$

$$\dot{g} + \frac{\ddot{g}}{\dot{g}} + 4\dot{A} = 0,$$

$$\ddot{\phi} + 4\dot{A}\dot{\phi} + \dot{g}\dot{\phi} - \frac{3}{8}e^{-g}\frac{dV}{d\phi} = 0.$$

Holographic Isotropic zero μ RG Flow

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Zero chemical potential

$$H_1 = 0$$

Non-zero temperature

$$Y \neq 0$$

Einstein-Dilaton E.O.M. are equivalent to HRG eqs:

Holographic Isotropic zero μ RG Flow

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Einstein-Dilaton E.O.M. are equivalent to HRG eqs:

$$\frac{dX}{d\phi} = -\frac{4}{3} (1 - X^2 + 1) \left(1 + \frac{3}{8} \frac{V'}{VX} \right)$$

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Holographic Isotropic zero μ RG Flow

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**Gursoy, Kiritsis, Mazzanti,
Nitti, arXiv:0812.0792**

Autonomous equations at finite T

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NEW VARIABLES

$$X = \frac{1}{3} \frac{\dot{\phi}}{\dot{A}(w)}, \quad Y = \frac{1}{4} \frac{\dot{g}}{\dot{A}}$$

and also

$$Z = \frac{1}{e^{\frac{2(16-9k^2)}{9k}\phi} + 1}, \quad V = C_1 \left(\frac{1-Z}{Z} \right)^{\frac{9k^2}{16-9k^2}} + C_2 \left(\frac{1-Z}{Z} \right)^{\frac{16}{16-9k^2}}$$

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Taking derivatives with respect to A from X, Y, Z we obtain

$$\frac{dZ}{dA} = \frac{2(9k^2 - 16)}{3k} (1 - Z) ZX,$$

$$\frac{dX}{dA} = (X^2 - 1 - Y) \frac{16C_2 + Z(9k^2C_1 - 16C_2) + 12kX(Z(C_1 - C_2) + C_2)}{3k((C_1 - C_2)Z + C_2)}$$

$$\frac{dY}{dA} = 4Y(X^2 - 1 - Y).$$

Fixed points

Fixed points

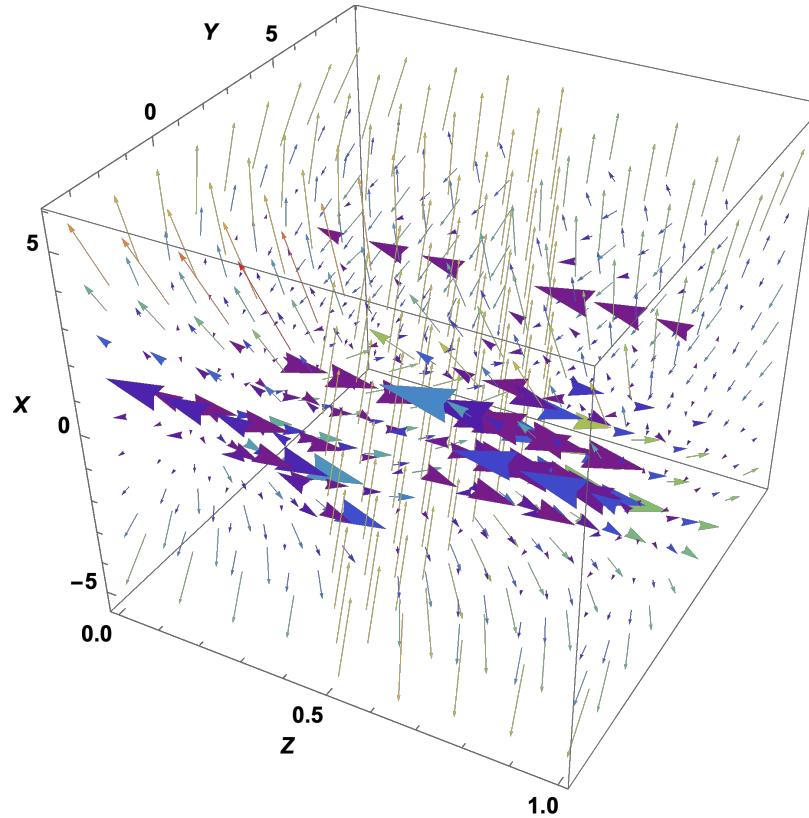


FIGURE: 3D phase portrait for the system at T with $k = 0.8$, the direction of arrows show the increasing energy scale (A)

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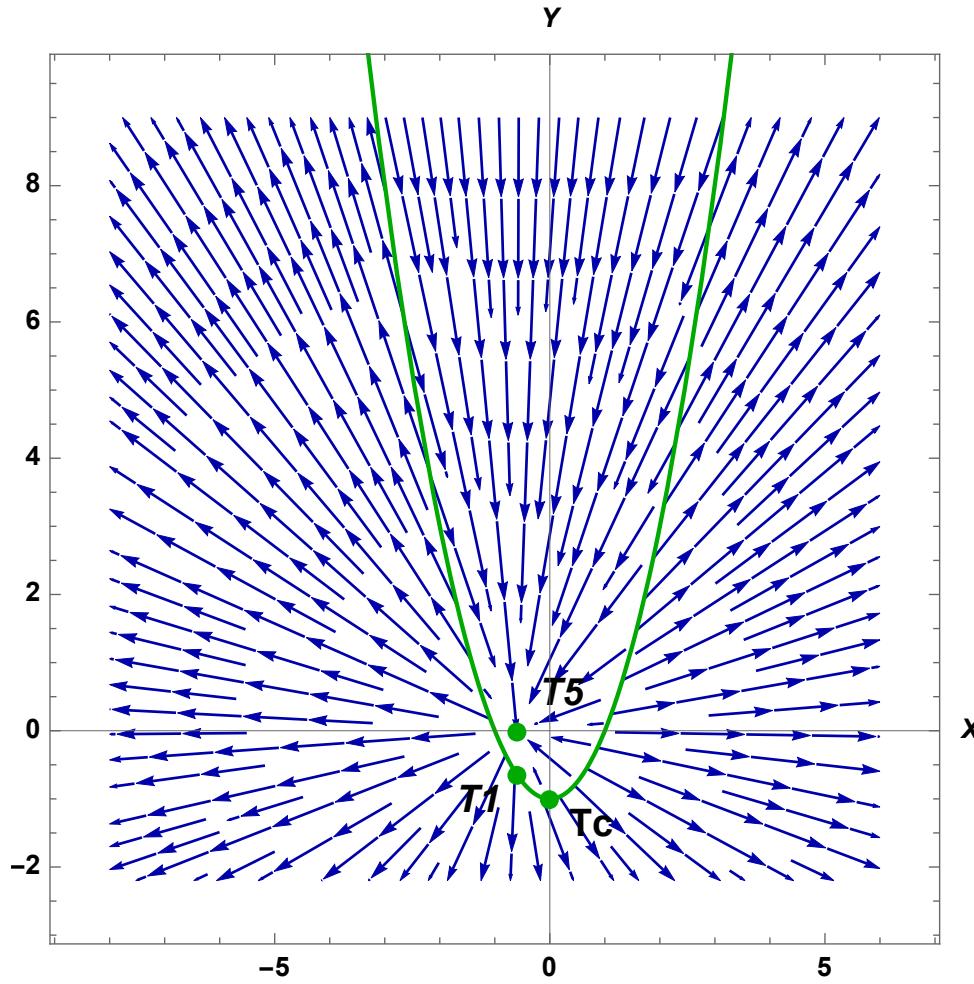


FIGURE: Phase portrait for the thermal system with $k = 0.8$, projection to $X - Y$ -plane with $Z = 1$

Backup

FIXED POINTS AT FINITE T

1. $X \in (-\infty; +\infty)$, $Y = X^2 - 1$, $Z = 1$. $\phi \rightarrow -\infty$, $V \rightarrow 0$. The eigenvalues

$$\lambda_1 = \frac{2}{3k}(16 - 9k^2)X, \quad \lambda_2 = 0, \quad \lambda_3 = 4(X^2 + \frac{3k}{2}X + 1).$$

Since $0 < k < 4/3$ $\lambda_1 < 0$ for $X < 0$ and $\lambda_1 > 0$ for $X > 0$, λ_3 is always positive for $X \in (-\infty, \infty)$. For negative X we have a saddle while for positive this fixed point is unstable node (**a saddle-node bifurcation**). For $X < 0$ the β -function is negative while for $X > 0$ it is positive.

$$\frac{dX}{dA} = (X^2 - 1 - Y)(3k + 4X),$$

2. $X \in (-\infty; +\infty)$, $Y = X^2 - 1$, $Z = 0$.

$$\lambda_1 = \frac{2}{3k}(9k^2 - 16)X, \lambda_2 = 0, \lambda_3 = 4 \left(X + \frac{\sqrt{16 - 9k^2} + 4}{3k} \right) \left(X - \frac{\sqrt{16 - 9k^2} - 4}{3k} \right).$$

$\lambda_1 > 0$ with $X < 0$, $\lambda_1 < 0$ with $X > 0$, $\lambda_3 < 0$ for $X \in \left(-\frac{\sqrt{16 - 9k^2} + 4}{3k}; \frac{\sqrt{16 - 9k^2} - 4}{3k} \right)$, while for $X \in \left(-\infty, -\frac{\sqrt{16 - 9k^2} + 4}{3k} \right) \cup \left(\frac{\sqrt{16 - 9k^2} - 4}{3k}, +\infty \right)$ $\lambda_3 > 0$. Choosing k and X

we can get **a)** both unstable and stable nodes, **b)** saddle fixed points.

HRGF for non-zero chemical potential and anisotropic metric

5-dim Background

5-dim Background

Einstein-dilaton-two-Maxwell

5-dim Background

Einstein-dilaton-two-Maxwell

$$S = \int \frac{d^5x}{16\pi G_5} \sqrt{-\det(g_{\mu\nu})} \left[R - \frac{f_1(\phi)}{4} F_{(1)}^2 - \frac{f_2(\phi)}{4} F_{(2)}^2 - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - V(\phi) \right]$$

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I.A, K.Rannu,
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$$\lambda = e^\phi$$

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$$E \sim B$$

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Holographic Anizotropic RenormGroup Flow

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Non-zero Aniz.: $H_2 \neq 0$

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I.A., K.Rannu

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Thank you for your attention